

Lecture 13 c Coupled oscillators

Consider a system of N atoms interacting via an N -body potential:

$$\hat{H} = \sum_{i=1}^N \frac{\vec{p}_i^2}{2m} + U(\{\vec{r}_i\})$$

At low temperatures, the configuration will be close to the minimum of the potential:

$$U_0 = \min \{ U(\{\vec{r}_i\}) \} = U(\{\vec{r}_i^{(0)}\})$$

Then we can Taylor expand

$$U(\{\vec{r}_i\}) \approx U_0 + \frac{1}{2} \sum_{i,j} \vec{x}_i \cdot \left. \frac{\partial^2 U}{\partial \vec{r}_i \partial \vec{r}_j} \right|_{\{\vec{r}_k^{(0)}\}} \cdot \vec{x}_j$$

The 1st-order term is zero.

Here $\vec{x}_i = \vec{r}_i - \vec{r}_i^{(0)}$.

This gives us a Hamiltonian
of many coupled oscillators

$N_{\text{modes}} = dN$, for $d=1, 2, 3$
spatial dimensions. We can

find the normal modes X_α :

$$\hat{H} = \sum_{\alpha=1}^{N_{\text{modes}}} \left(\frac{p_\alpha^2}{2m} + \frac{m\omega_\alpha^2}{2} X_\alpha^2 \right),$$

where the transformation

$$\{ \vec{x}_i \} \rightarrow \{ X_\alpha \} \quad \text{is an}$$

orthogonal transformation.

Quantizing the problem, we have

$$\hat{H} = \sum_{\alpha=1}^{N_{\text{modes}}} \hbar\omega_\alpha \left(a_\alpha^\dagger a_\alpha + \frac{1}{2} \right),$$

where $[a_\alpha, a_\beta^\dagger] = \delta_{\alpha\beta}$.

3

The mean-square displacement of an atom from its equilibrium position is

$$\langle \bar{x}_n^2 \rangle = \frac{1}{N} \sum_{i=1}^N \langle \bar{x}_i^2 \rangle = \frac{1}{N} \sum_{\alpha=1}^{N_{\text{modes}}} \langle x_\alpha^2 \rangle.$$

The last equality holds because the transformation is orthogonal.

$$x_\alpha = \sqrt{\frac{\hbar}{2m\omega_\alpha}} (a_\alpha^\dagger + a_\alpha)$$

$$\langle x_\alpha^2 \rangle = \frac{\hbar}{2m\omega_\alpha} \langle (a_\alpha^\dagger)^2 + (a_\alpha)^2 + a_\alpha^\dagger a_\alpha + a_\alpha a_\alpha^\dagger \rangle$$

In an energy eigenstate

$$\begin{aligned} \langle n_\alpha | (a_\alpha^\dagger)^2 + (a_\alpha)^2 + a_\alpha^\dagger a_\alpha + a_\alpha a_\alpha^\dagger | n_\alpha \rangle \\ = 2n_\alpha + 1 \end{aligned}$$

In equilibrium at temperature T , (4)

$$P(n_\alpha) = \frac{e^{-\beta E_{n_\alpha}}}{Z_\alpha} = \frac{e^{-\beta \hbar \omega_\alpha (n_\alpha + \frac{1}{2})}}{Z_\alpha}$$

$$Z_\alpha = \sum_{n_\alpha=0}^{\infty} e^{-\beta \hbar \omega_\alpha (n_\alpha + \frac{1}{2})} \quad (\text{Here } \beta = \frac{1}{k_B T})$$

$$= \frac{e^{-\beta \hbar \omega_\alpha / 2}}{1 - e^{-\beta \hbar \omega_\alpha}}$$

$$\langle E_\alpha \rangle = - \frac{\partial \ln Z_\alpha}{\partial \beta} = \hbar \omega_\alpha \left(\langle n_\alpha \rangle + \frac{1}{2} \right)$$

$$= \hbar \omega_\alpha \left(\frac{1}{e^{\beta \hbar \omega_\alpha} - 1} + \frac{1}{2} \right)$$

$$\langle n \rangle = \frac{1}{e^{\beta \hbar \omega} - 1}$$

Planck
distribution

$$\langle X_d^2 \rangle = \frac{\hbar}{2m\omega_d} \left(\frac{1}{e^{\beta\hbar\omega_d} - 1} + \frac{1}{2} \right)$$

5

$$\langle \vec{X}_n^2 \rangle = \frac{1}{N} \sum_{d=1}^{N_{\text{modes}}} \frac{\hbar}{2m\omega_d} \left(\frac{1}{e^{\beta\hbar\omega_d} - 1} + \frac{1}{2} \right)$$

Debye model

$$\omega_s(\vec{k}) = v|\vec{k}|, \quad s = 1, \dots, d$$

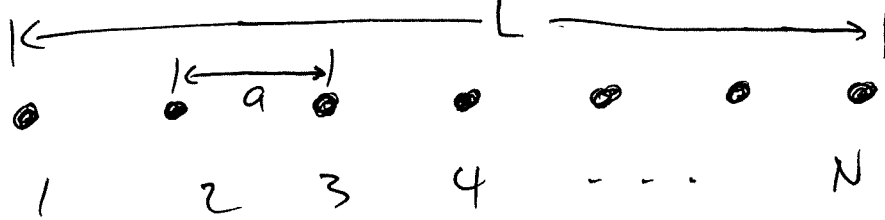
$$v = \sqrt{B/\rho} = \text{speed of sound in materials}$$

i) 1 dimension

$$\langle X_n^2 \rangle = \frac{1}{N} \int \frac{dx dp}{h} \frac{\hbar^2}{2mV|p|} \left(\frac{1}{e^{\beta V|p|} - 1} + \frac{1}{2} \right)$$

$$= \frac{L}{N} \frac{\hbar}{4\pi V} \int_{-\frac{\hbar\pi}{a}}^{\frac{\hbar\pi}{a}} dp \frac{1}{|p|} \left(\frac{1}{e^{\beta V|p|} - 1} + \frac{1}{2} \right)$$

$\rightarrow \infty$ even at $T=0!$



6

Quantum fluctuations diverge!

Even though the mean positions of the atoms are arranged as shown above, the fluctuations about these mean positions are unbounded (for a very long system), indicating that such a system is a fluid, not a solid.

Harmonic fluids are ubiquitous in 1d interacting systems -

For the 2d case, one can

7

show that

$$\langle \vec{x}_n^2 \rangle \Big|_{T=0} = \text{finite} \quad \text{but}$$

$$\langle \vec{x}_n^2 \rangle \Big|_{T>0} \rightarrow \infty$$

A 2d crystal is not destroyed by quantum fluctuations, but is destroyed by thermal fluctuations at any finite temperature for a very large system. In 3d, everything is bounded for $T \geq 0$.

\Rightarrow Mermin-Wagner theorem